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ITALY

A new Proton Recoil Telescope for the n_TOF@CERN facility

The Motivation

The role of the $^{235}\text{U}(n,f)$ reaction



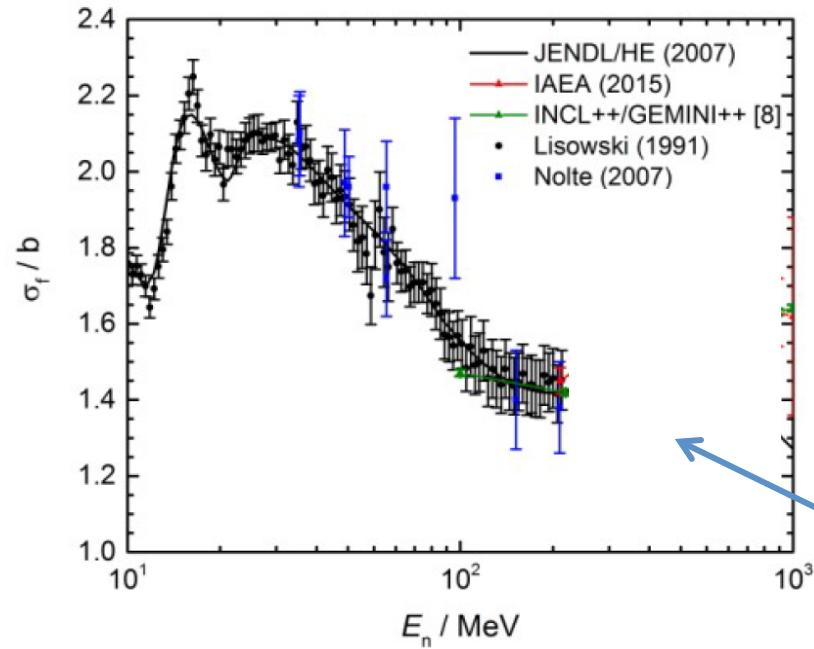
Neutron Cross Section Standards are essential for measurements and evaluations of other neutron cross sections

The $^{235}\text{U}(n,f)$ cross section is one of the most important standards for neutron fluence measurements and for several applications, such as accelerator-driven nuclear systems, biological effectiveness of high-energy neutrons, etc..

$^{235}\text{U}(n,f)$ is a standard at

- **Thermal energy: 0.025 eV**
- **Energy range: 0.15MeV ÷ 200MeV.**

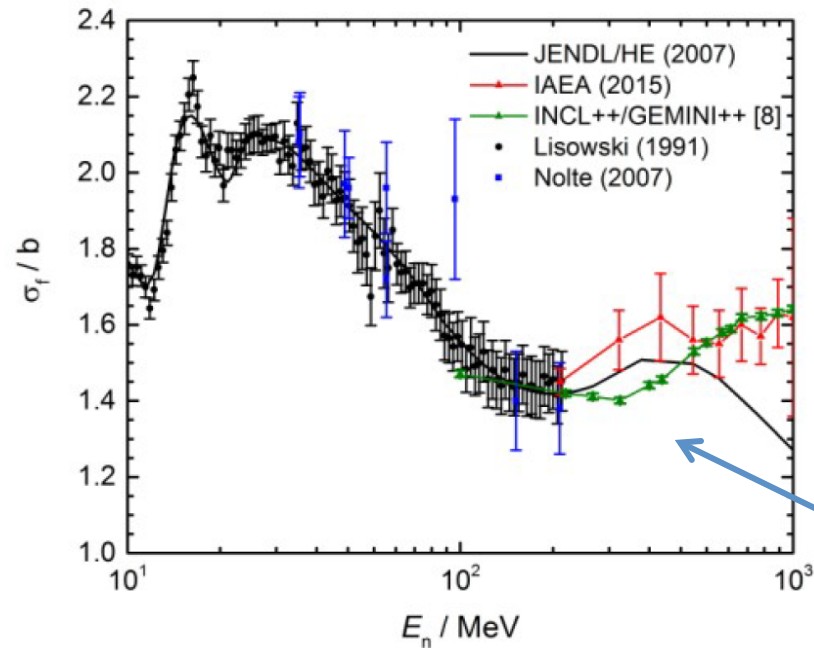
Status of the $^{235}\text{U}(n,f)$ cross section



Two sets of experimental data in the range 20 MeV – 200 MeV

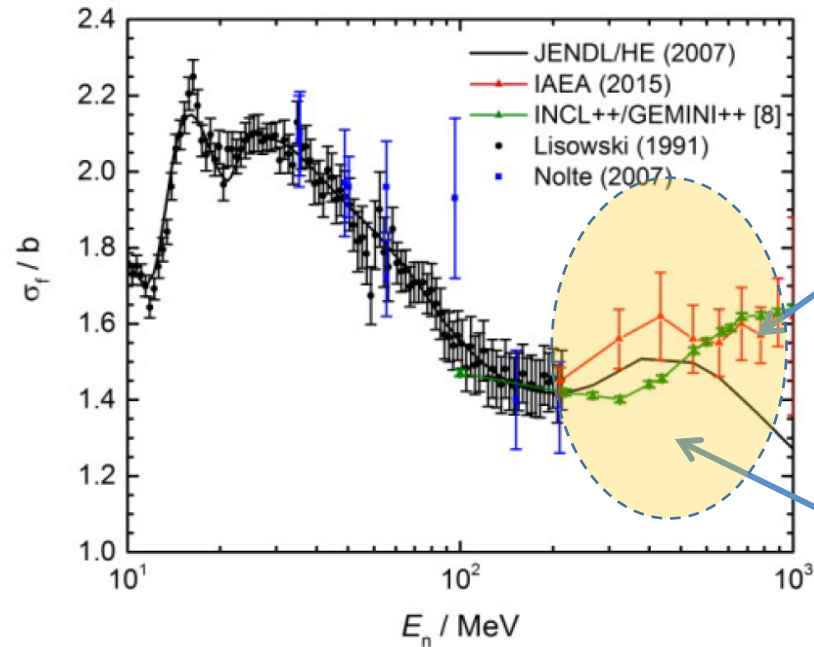
No experimental data above 200MeV. Only theoretical models

Status of the $^{235}\text{U}(n,f)$ cross section



Two sets of experimental data in the range 20 MeV – 200 MeV

No experimental data above 200MeV. Only theoretical models



The theoretical cross section above 200MeV may be substantially different, depending on the model used for the calculations.

No experimental data above 200MeV. Only theoretical models

Two sets of experimental data in the range 20 MeV – 200 MeV

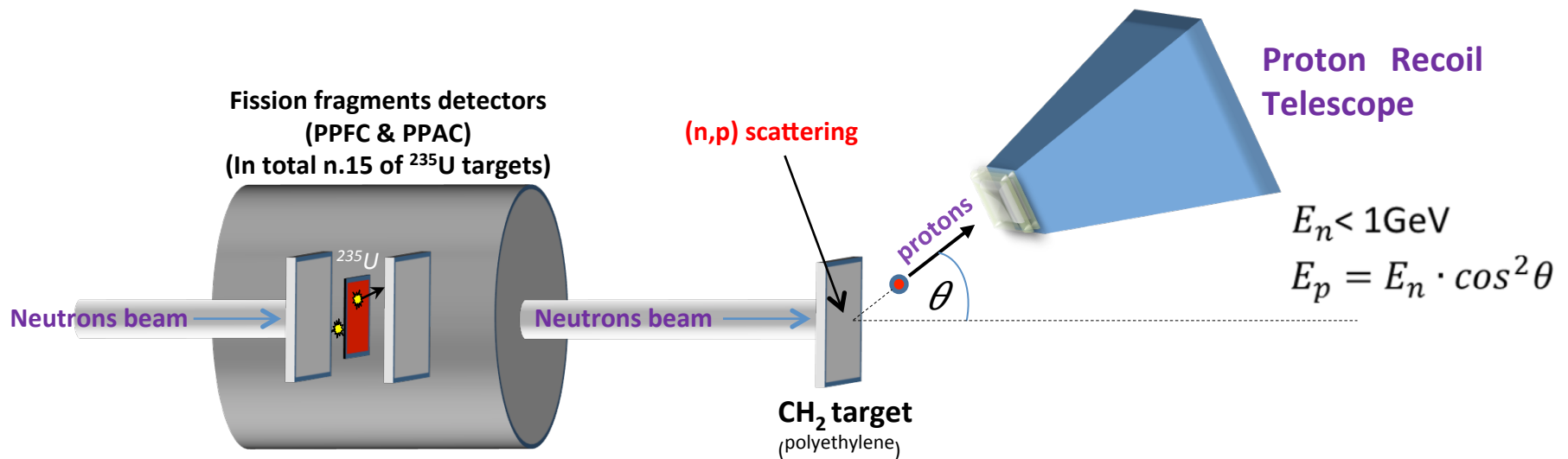
**The International Atomic Energy Agency (IAEA) strongly requests new data for the $^{235}\text{U}(n,f)$ cross section up to 1GeV, with the aim to improve the uncertainty within 5%.
New measurements above 20 MeV are needed!**

The Technique: Two simultaneous measurements with a suitable neutron beam.

- $^{235}\text{U}(n,f)$ reaction vs neutron energy
- $\text{H}(n,n)\text{H}$ reaction vs neutron energy

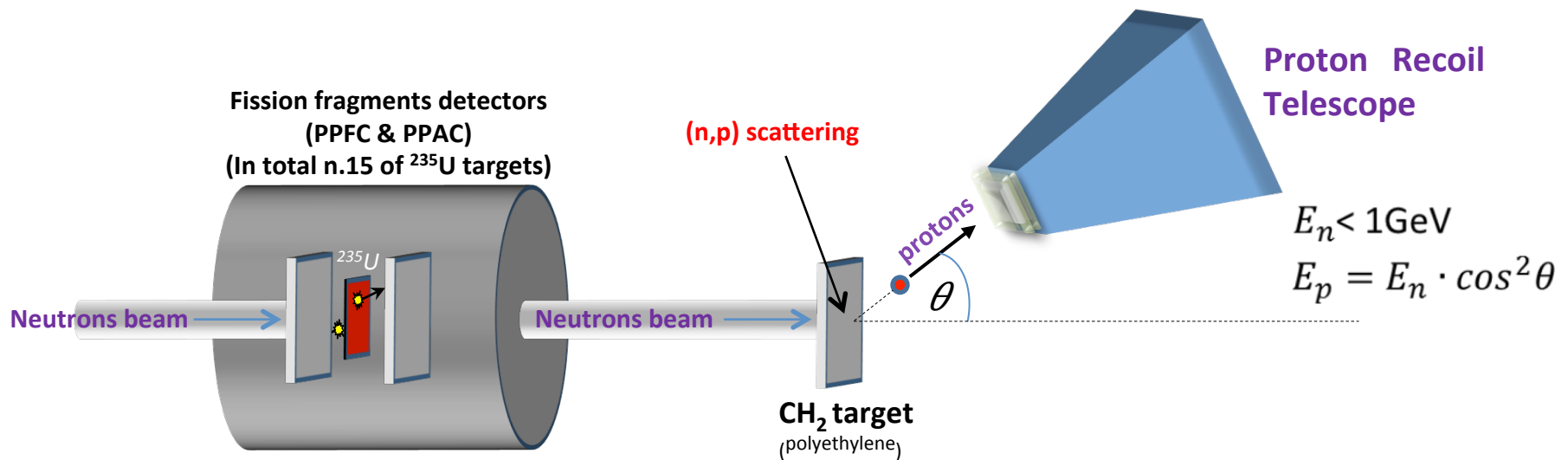
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The Technique: Two simultaneous measurements with a suitable neutron beam.

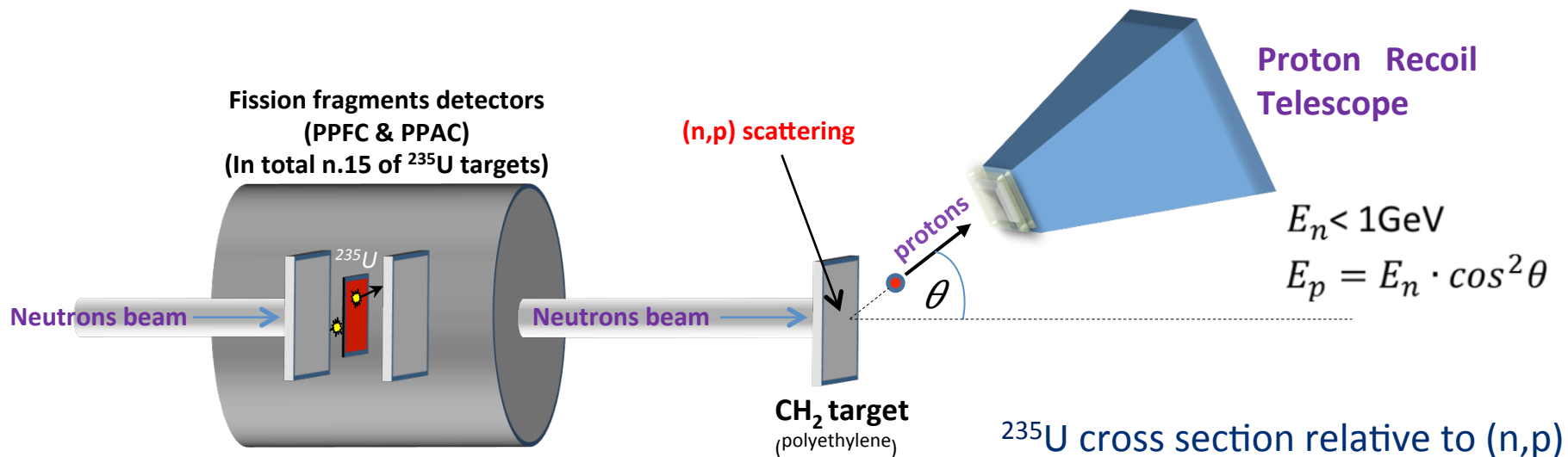
- $^{235}\text{U}(n,f)$ reaction vs neutron energy
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- The (n,p) scattering is used as a reference reaction for the $^{235}\text{U}(n,f)$ cross section
- The Proton Recoil Telescope (PRT) detects recoil protons emitted by the target, to measure the neutron flux vs. neutron energy

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^{235}U cross section relative to (n,p)

$$\frac{\sigma_{^{235}\text{U}(n,f)}}{(d\sigma_{np}/d\Omega)} = \frac{n_H \varepsilon_p \Omega_{\text{geo}} N_{\text{FF}}}{n_U \varepsilon_{\text{FF}} N_p}$$

N: detected events
 n: sample areal density
 Ω : PRT solid angle
 ε : efficiency
 : cross section

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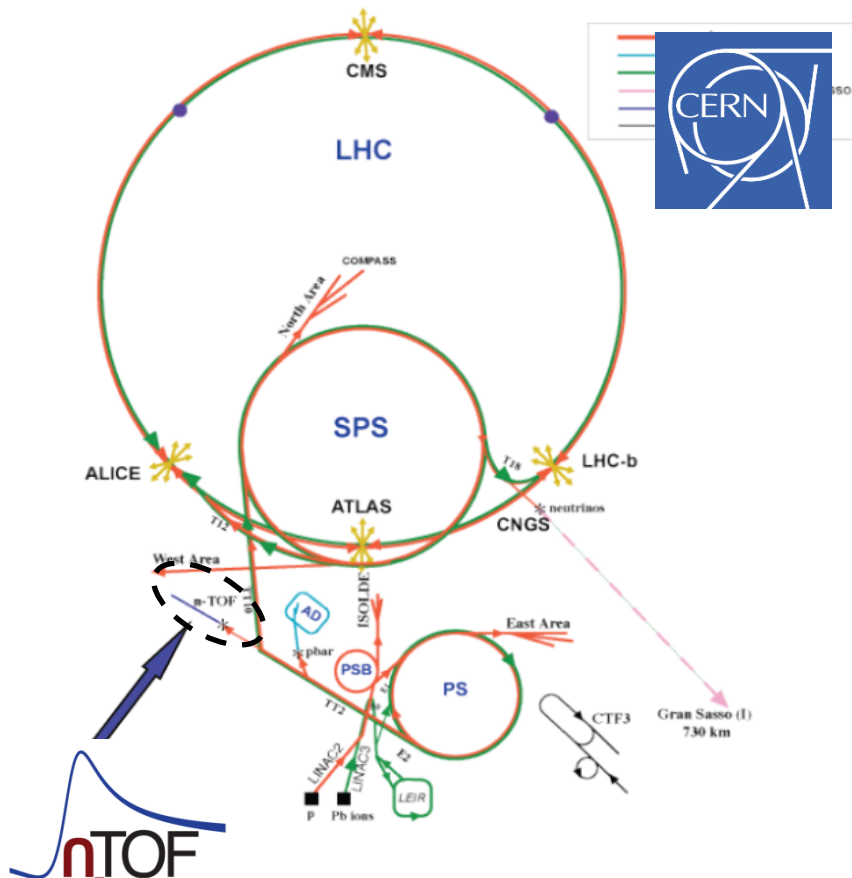
Where to measure the $^{235}\text{U}(n,f)$ cross section

A neutron beam with good energy resolution in a wide energy range is needed.



n_TOF@CERN

Neutron_Time-Of-Flight Facility

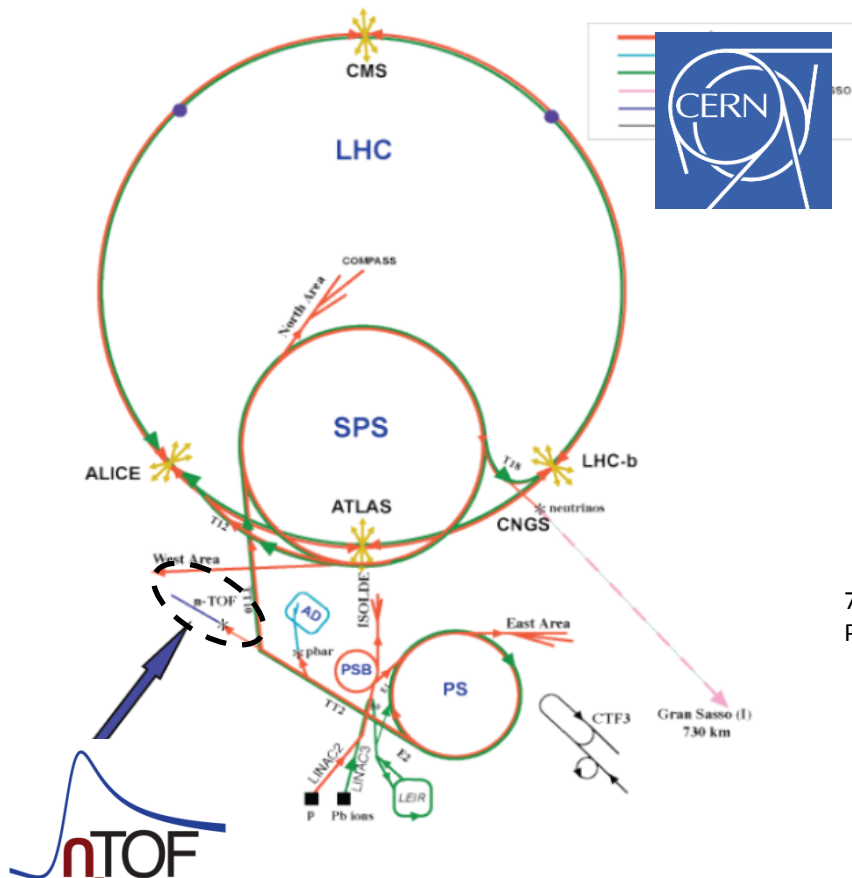


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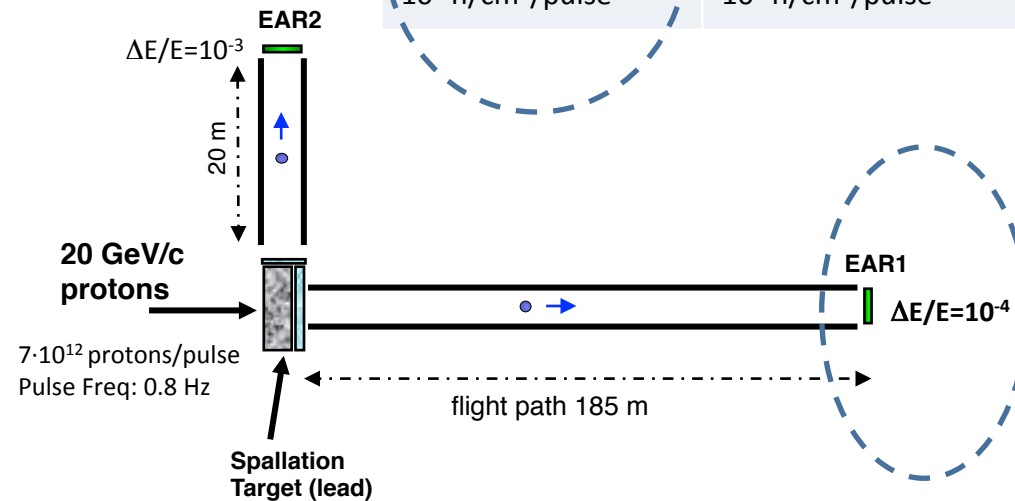
n_TOF@CERN

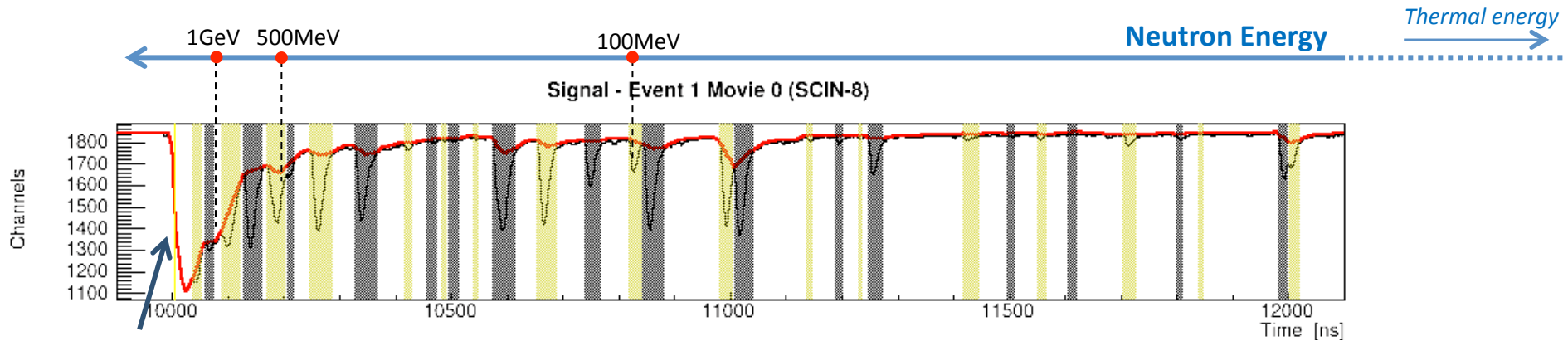
Neutron_Time-Of-Flight Facility



Two experimental areas

| EAR-1 | EAR-2 |
|--|--|
| $25 \text{ meV} < E_n < 1 \text{ GeV}$ | $25 \text{ meV} < E_n < 300 \text{ MeV}$ |
| $10^5 \text{ n/cm}^2/\text{pulse}$ | $10^6 \text{ n/cm}^2/\text{pulse}$ |



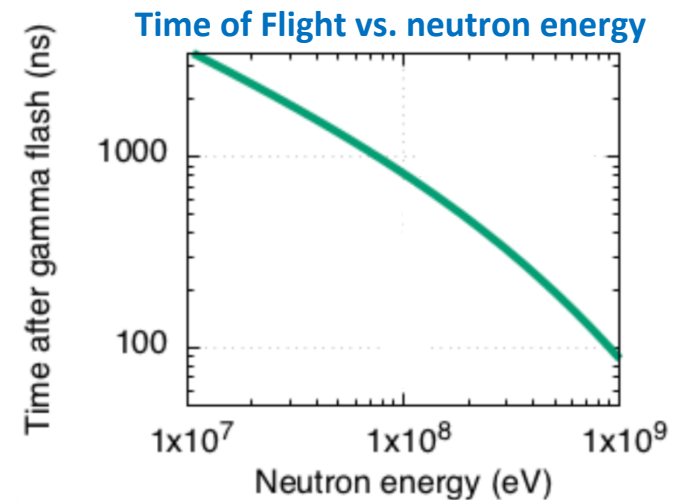


γ - Flash

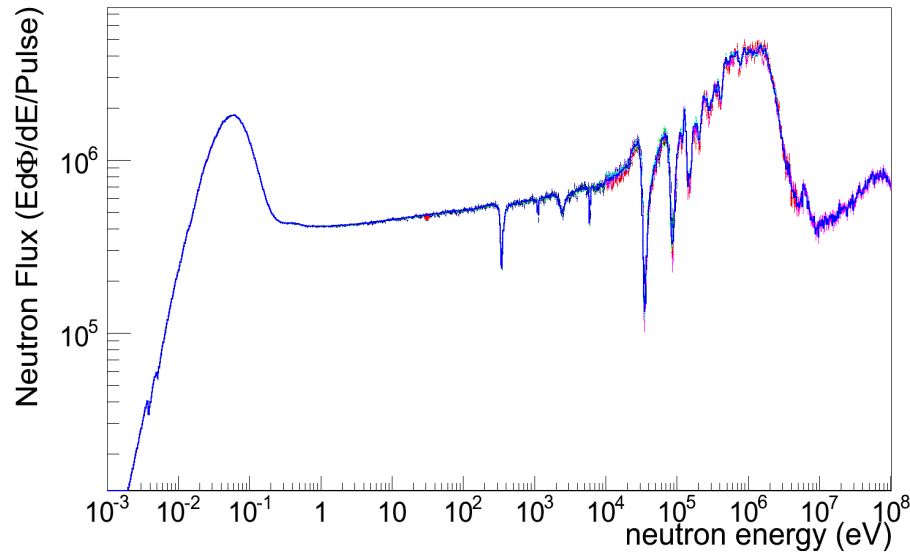
A typical time distribution of neutron - induced events in EAR1.

By subtracting the baseline the clean signals are extracted

- The γ -Flash is produced by the interaction of the protons pulse in the spallation target.
- It is the time reference to measure the time of flight of the incoming neutrons and hence their energy.
- Shorter the γ -Flash pulse width, higher the maximum kinetic neutron energy that can be detected

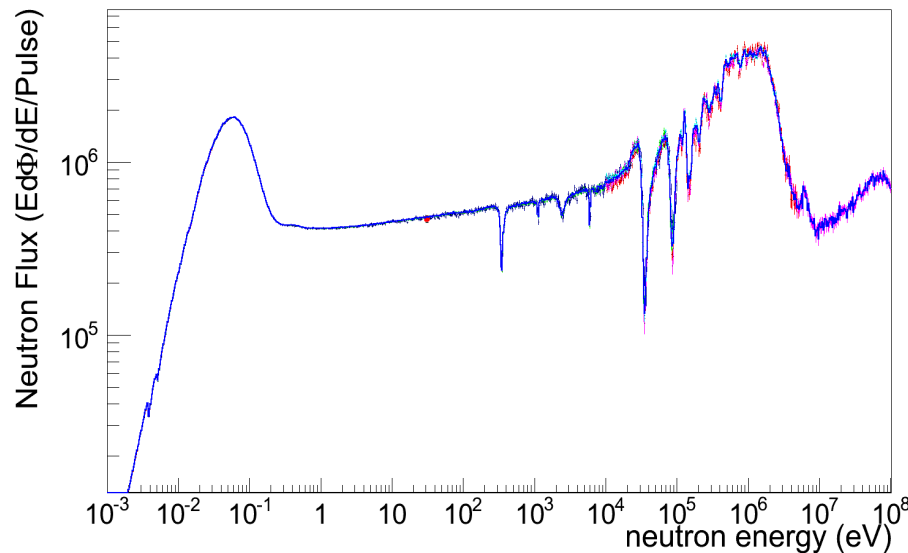


Energy distribution of the neutrons flux



n_TOF@CERN facility has all the requirements to measure the $^{235}\text{U}(n,f)$ cross section normalized to the (n,p) scattering.

Energy distribution of the neutrons flux



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The Proton Recoil Telescope

- To be used in a wide energy range: 10MeV - 1GeV
- $\Delta E-E$ particle identification to separate the recoil protons from the other light particles
- Working in coincidence mode to reject the background events
- Fast response to gamma γ -Flash

The Proton Recoil Telescope

Two different telescopes to be developed at n_TOF

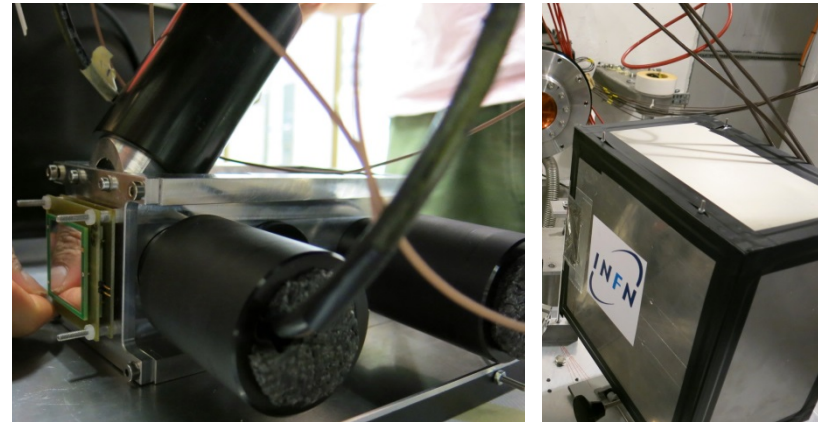
PTB - Germany

INFN - Italy

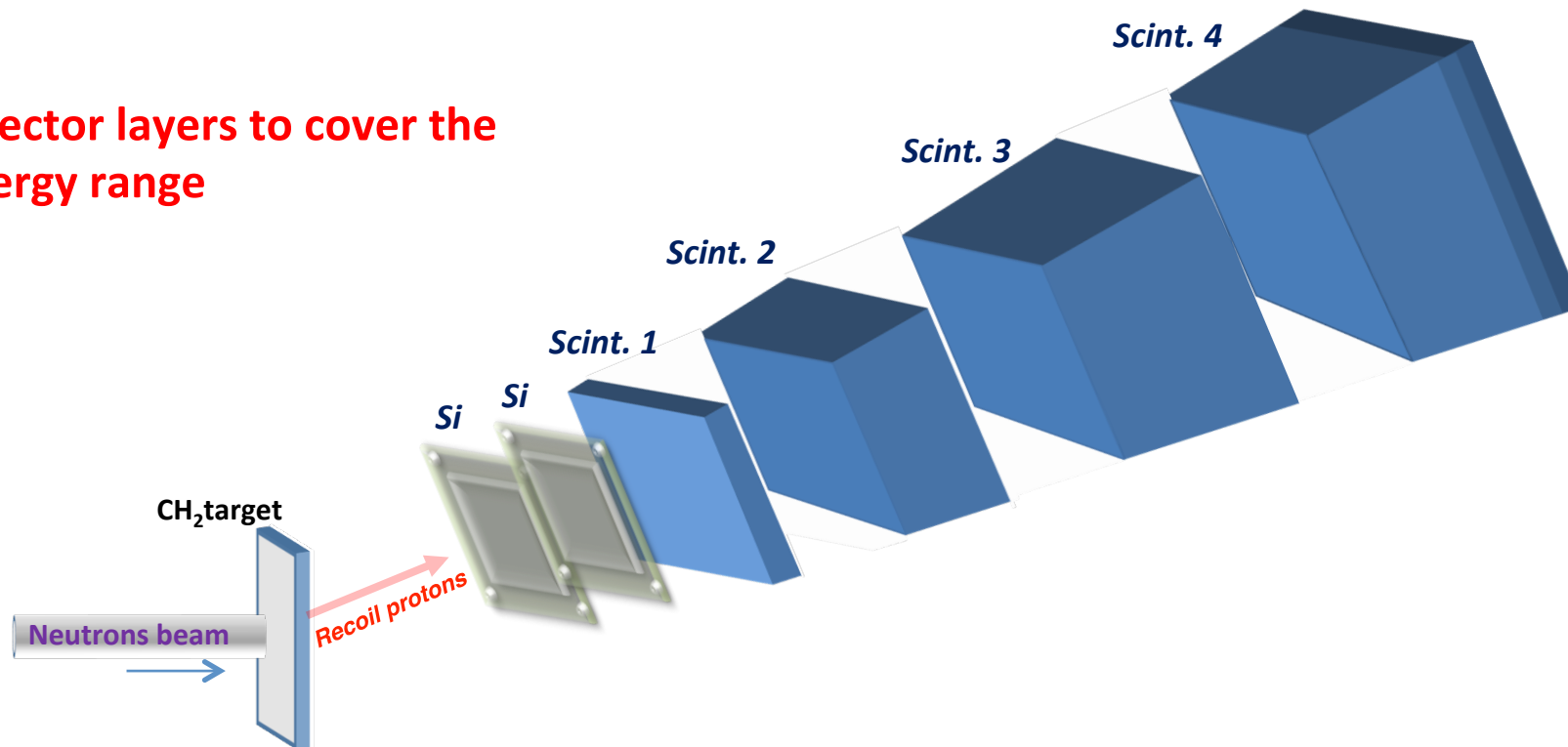
 **PTB** Physikalisch-Technische Bundesanstalt
Braunschweig und Berlin
Nationales Metrologieinstitut



 **INFN**
Istituto Nazionale
di Fisica Nucleare

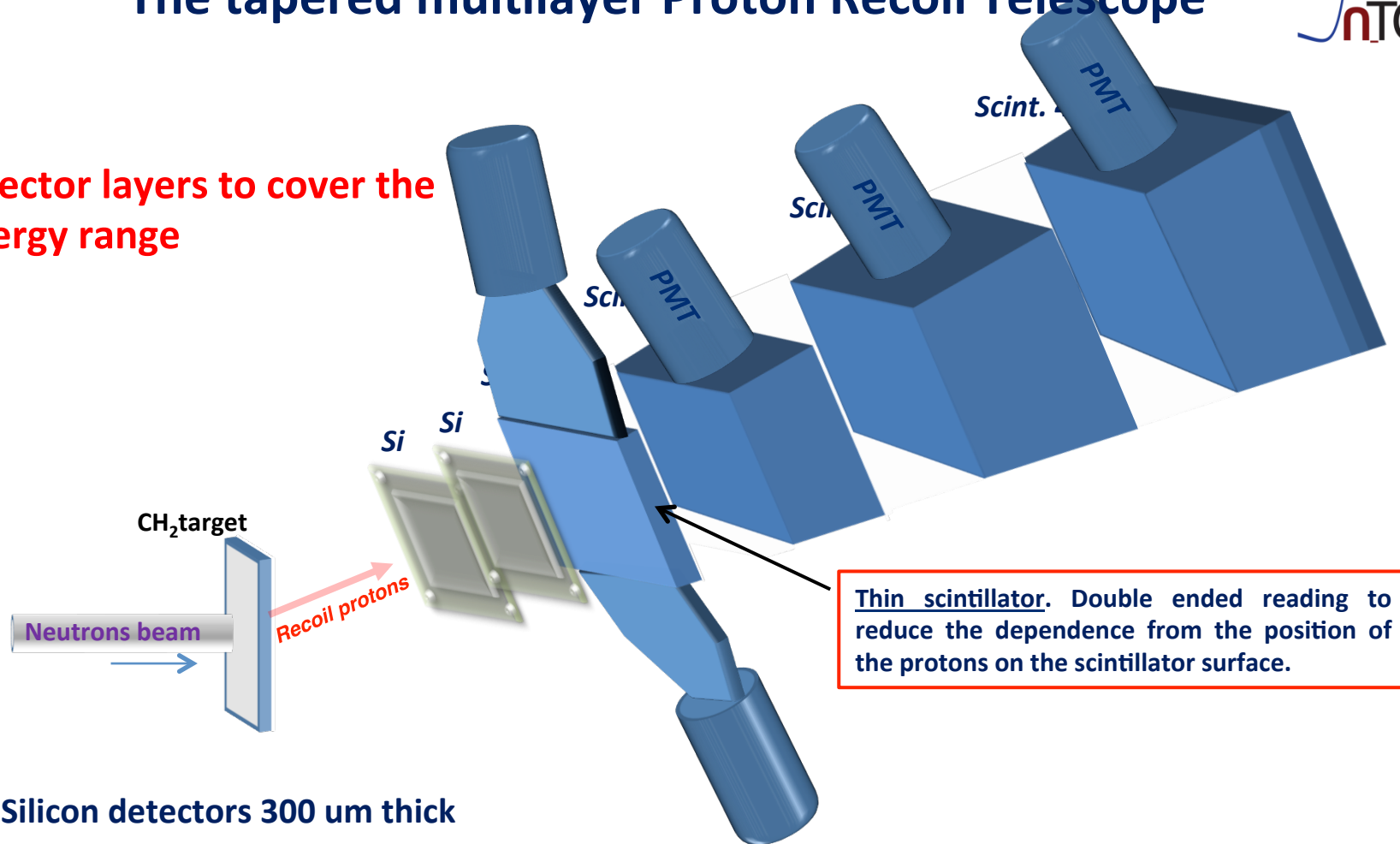


Six detector layers to cover the full energy range



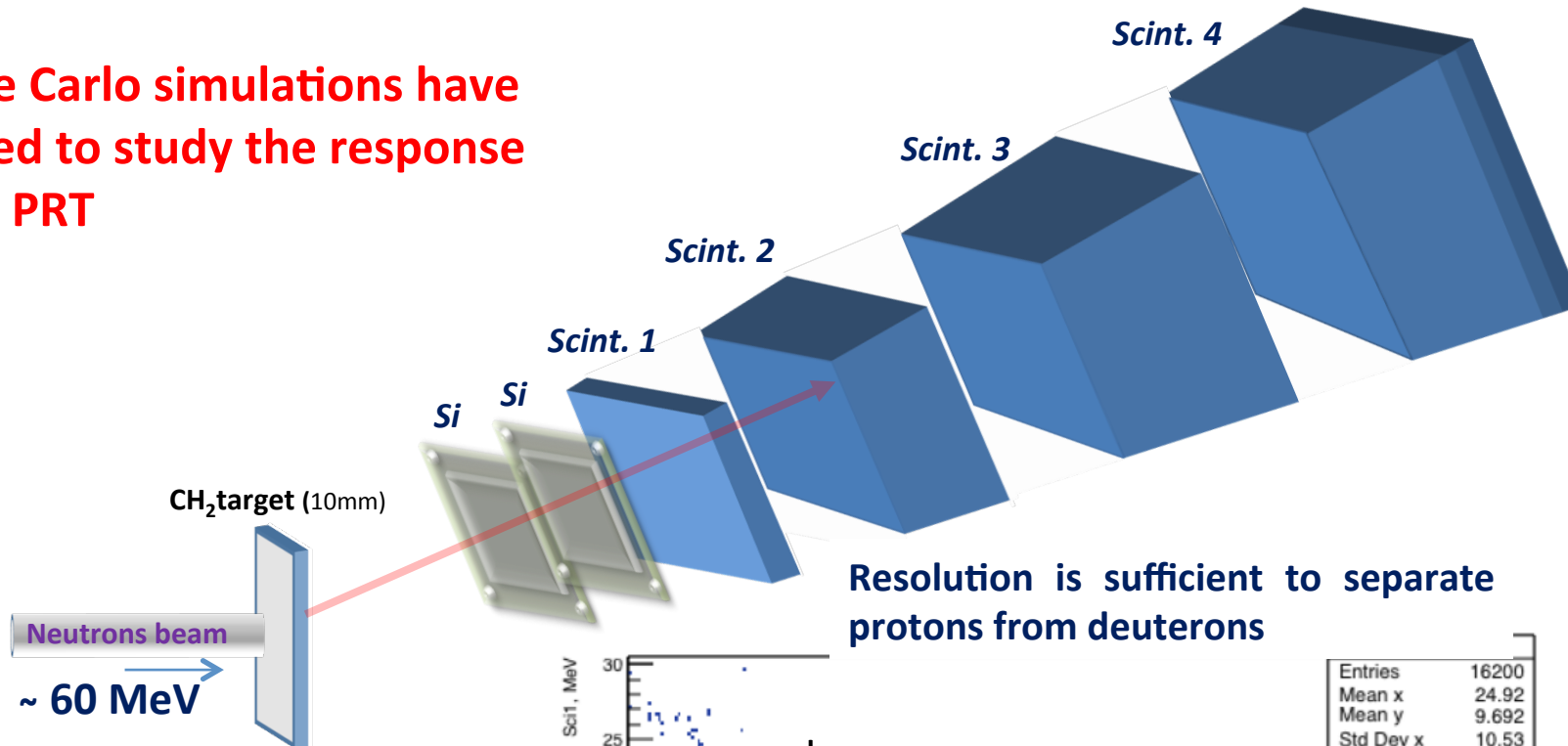
- 2 Silicon detectors 300 um thick
- 4 plastic scintillators BC408 ($\tau = 2.1$ ns) 0.5cm, 3cm, 6cm, 6cm
- Fast PMT (Hamamatsu R1924A)
- CH₂ target (*polyethylene*): 2mm – 10mm

Six detector layers to cover the full energy range

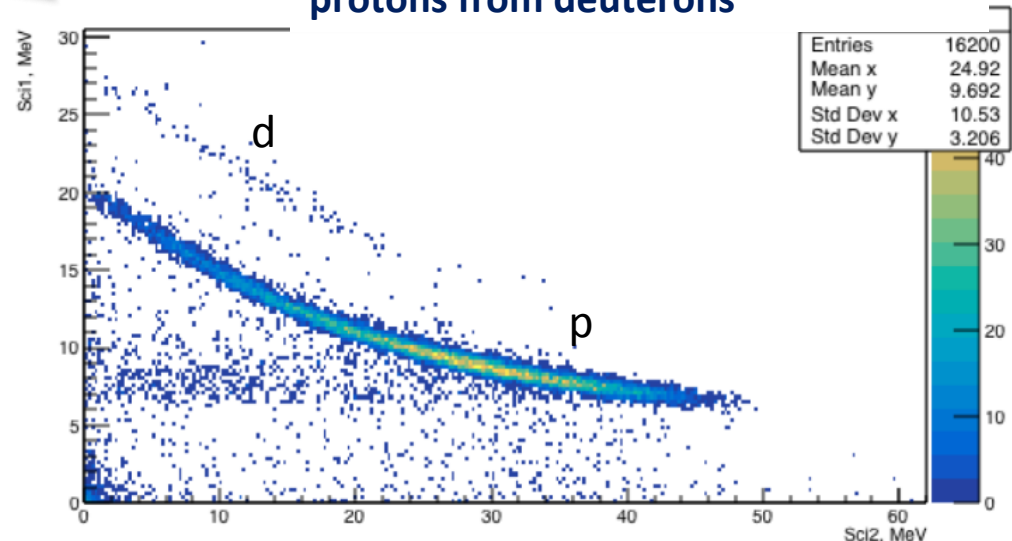


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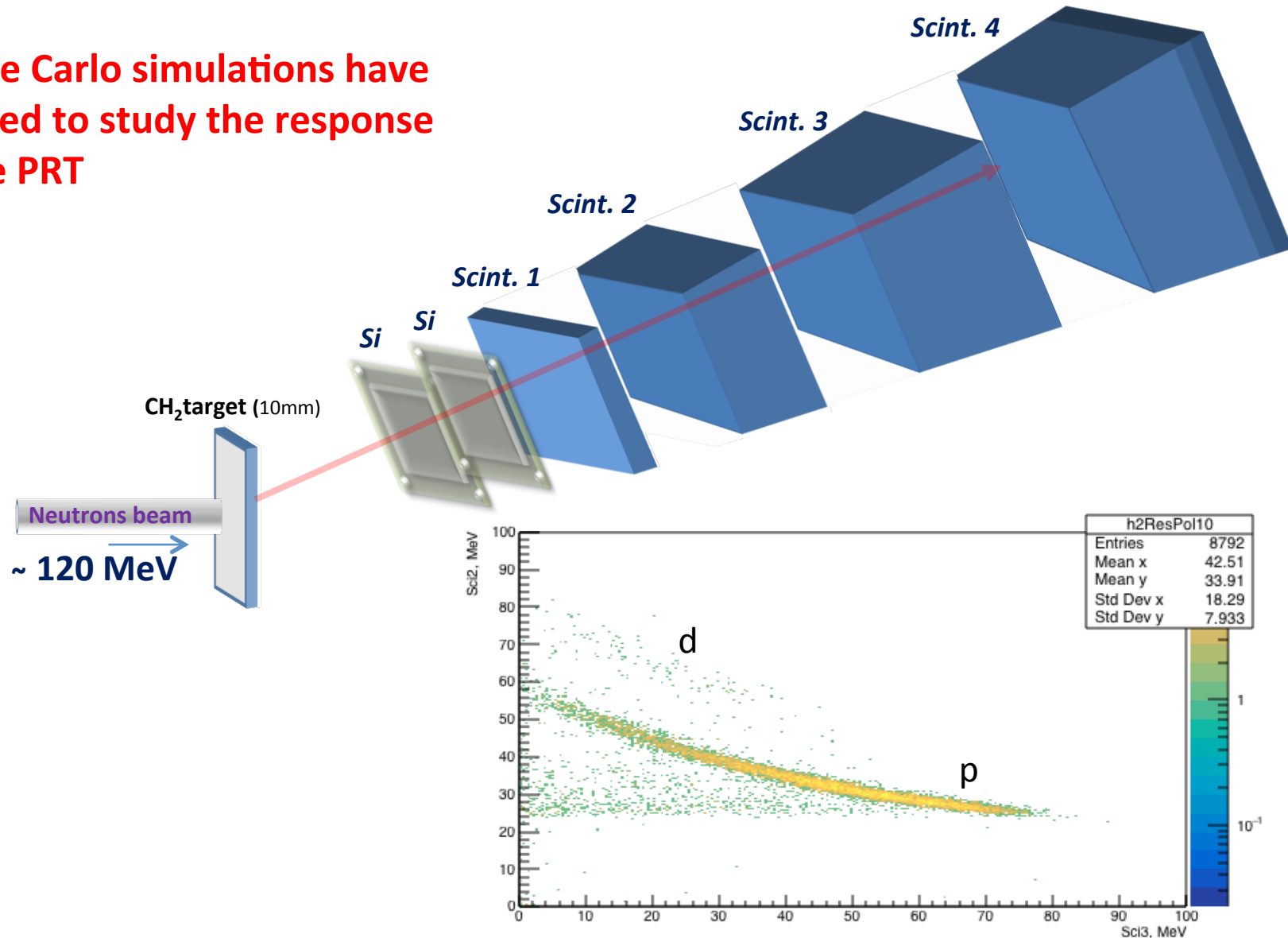
Monte Carlo simulations have allowed to study the response of the PRT



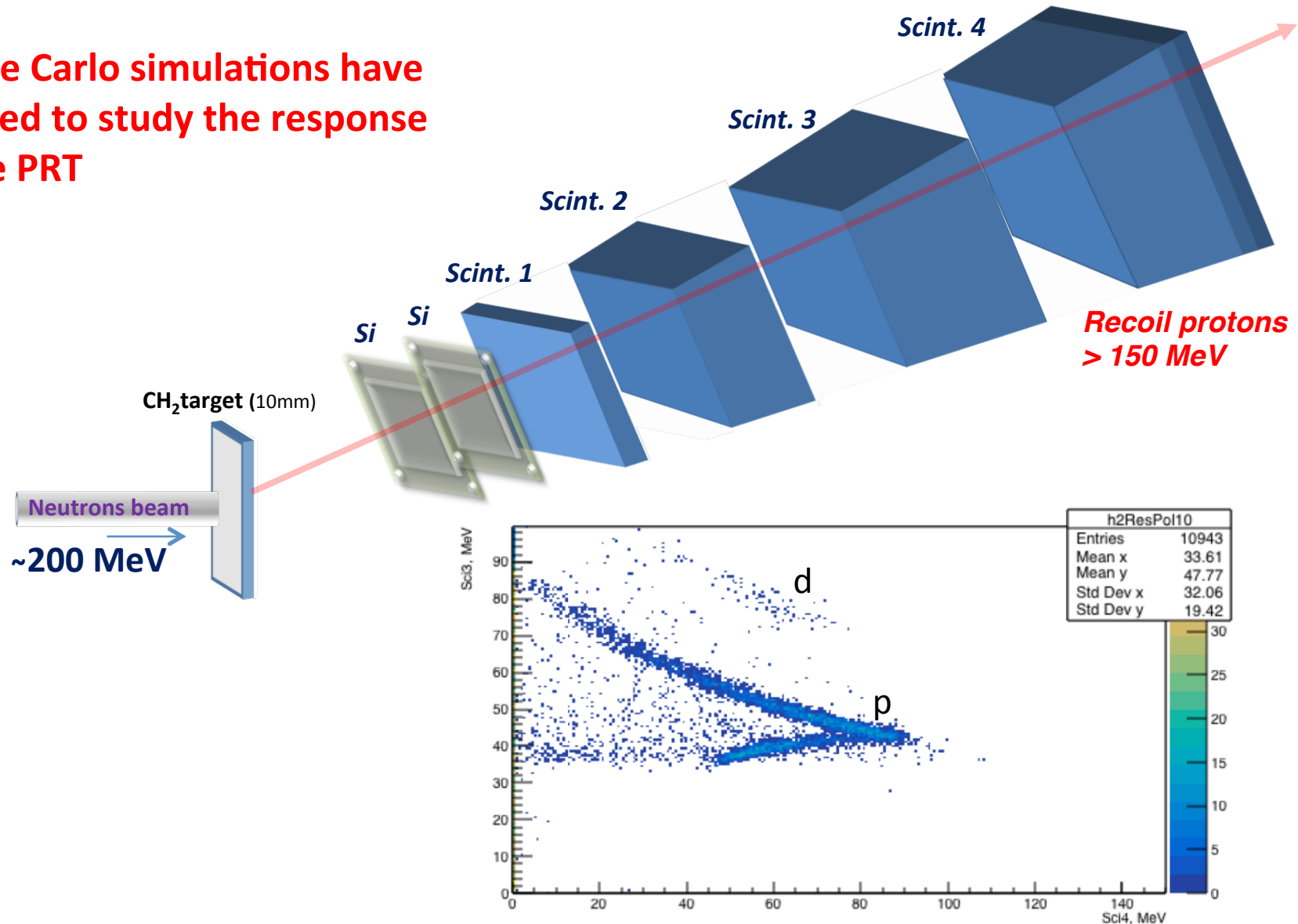
Resolution is sufficient to separate protons from deuterons



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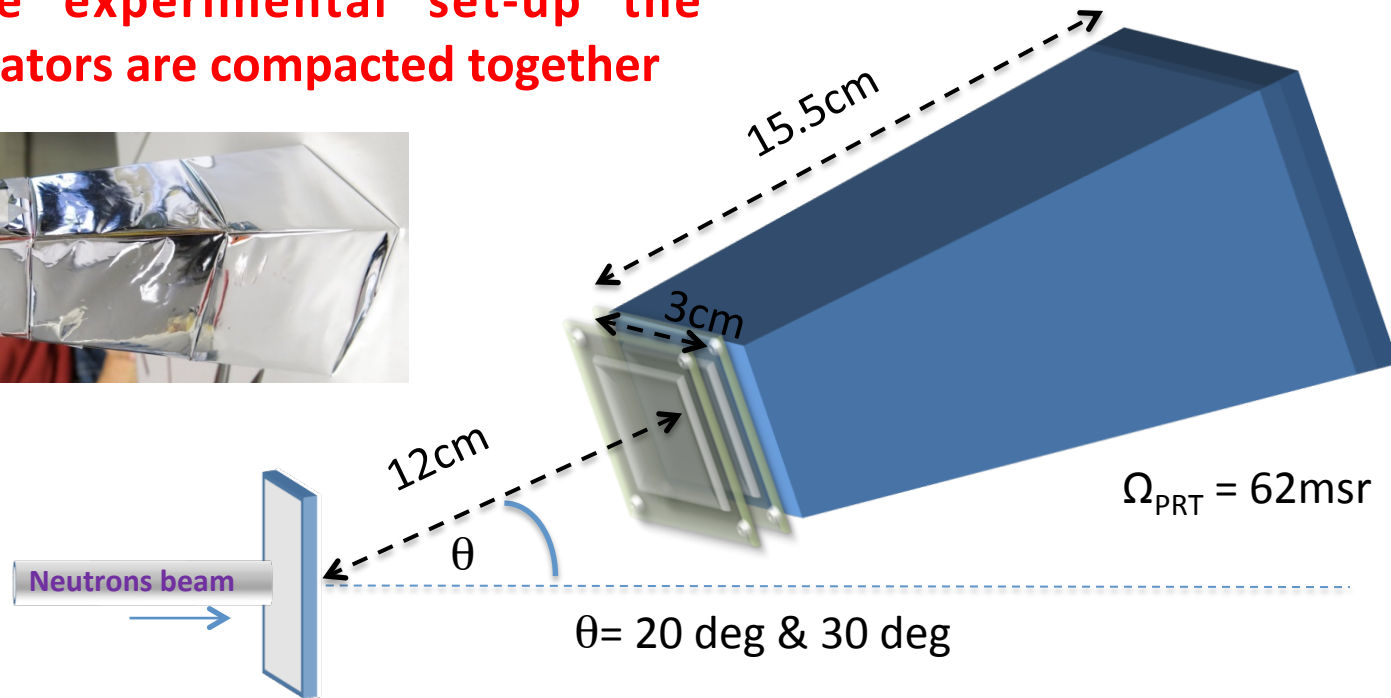
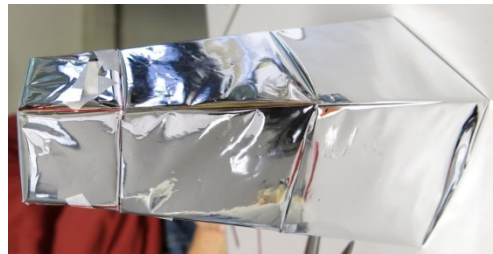


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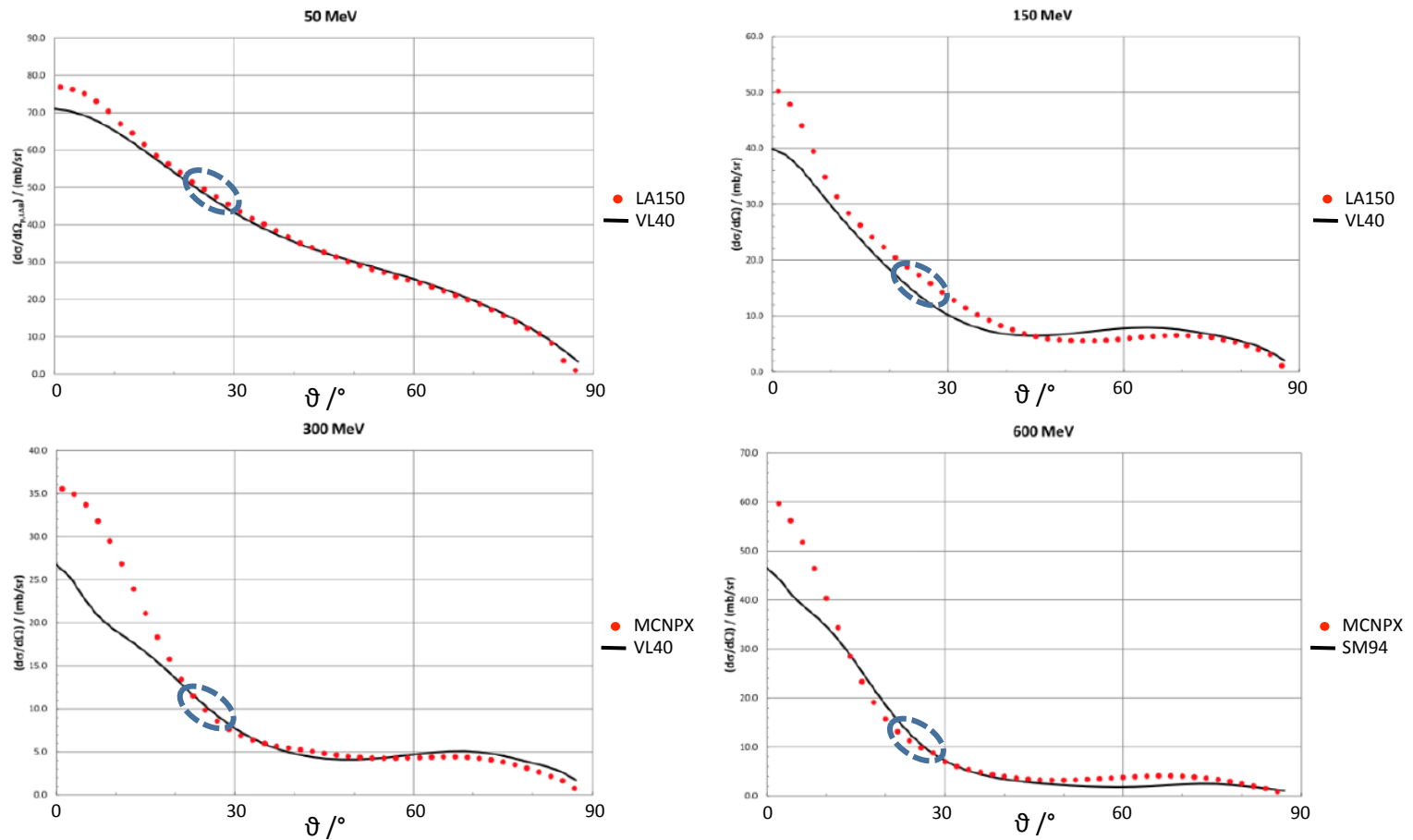
How to place the PRT

In the experimental set-up the scintillators are compacted together



The aim is to reach a statistical uncertainty around 2% within neutron energy regions of 5% relative width.

(n,p) cross section vs. angle of recoil protons

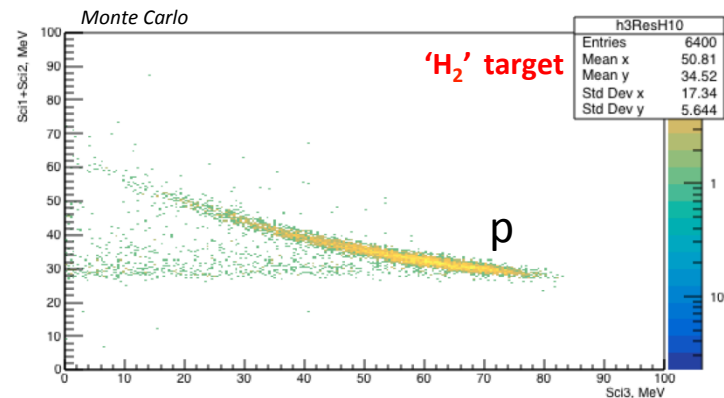
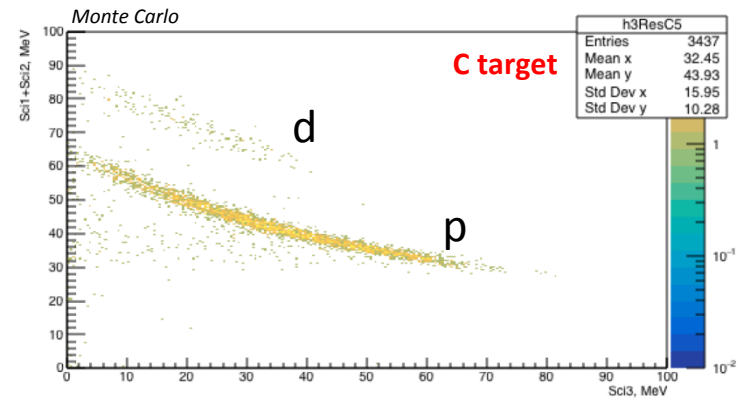
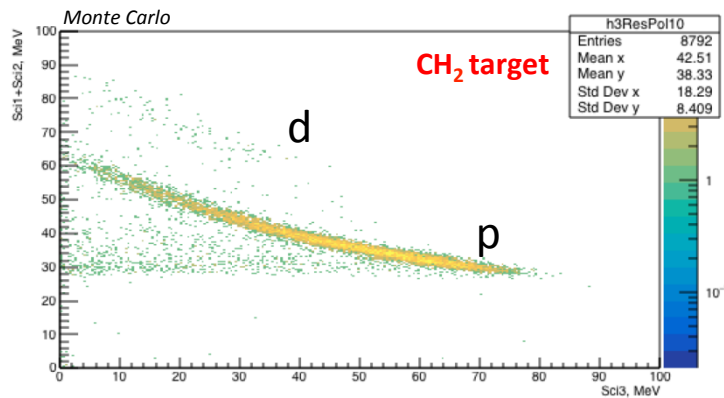
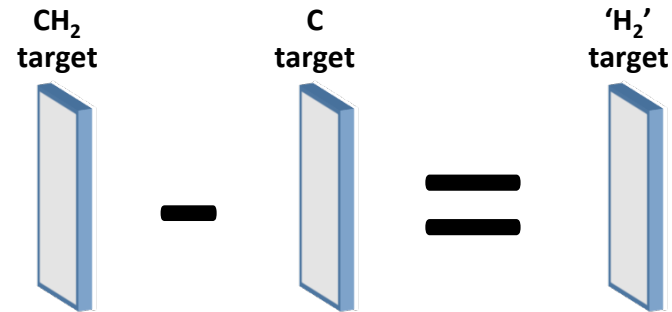


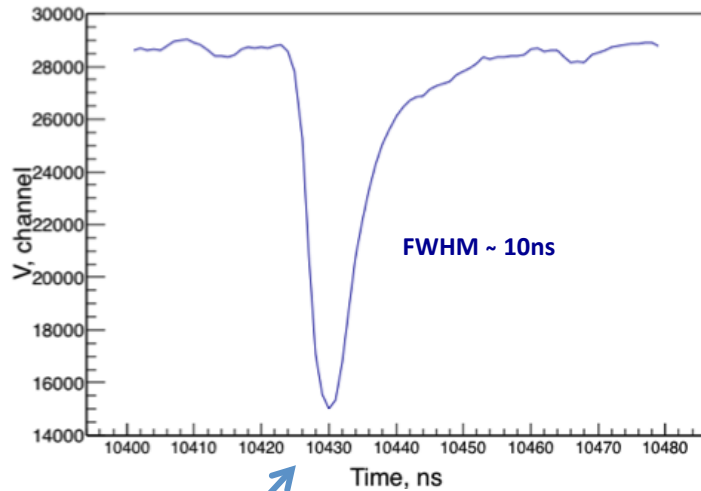
20 – 30 deg

Theoretical team is working to quantify the uncertainties

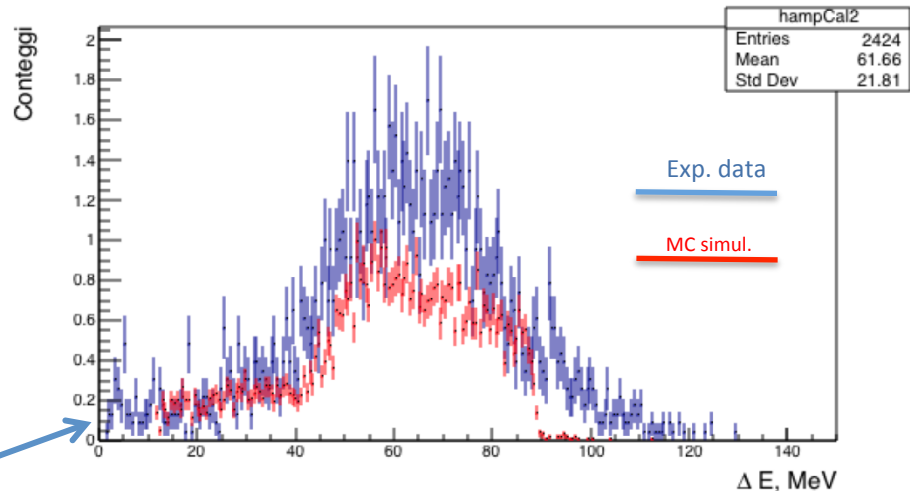
Select only the events due to hydrogen atoms in the CH₂ target

Measurements with graphite targets to subtract the contribution of the competing reaction ¹²C(n,x) in the CH₂ (polyethylene) target



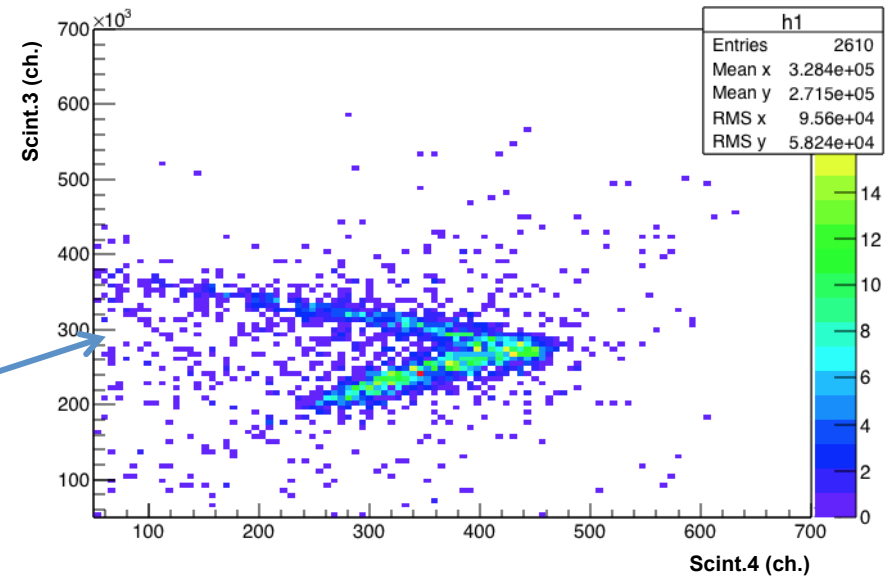


Typical pulse
PMT+scintillator



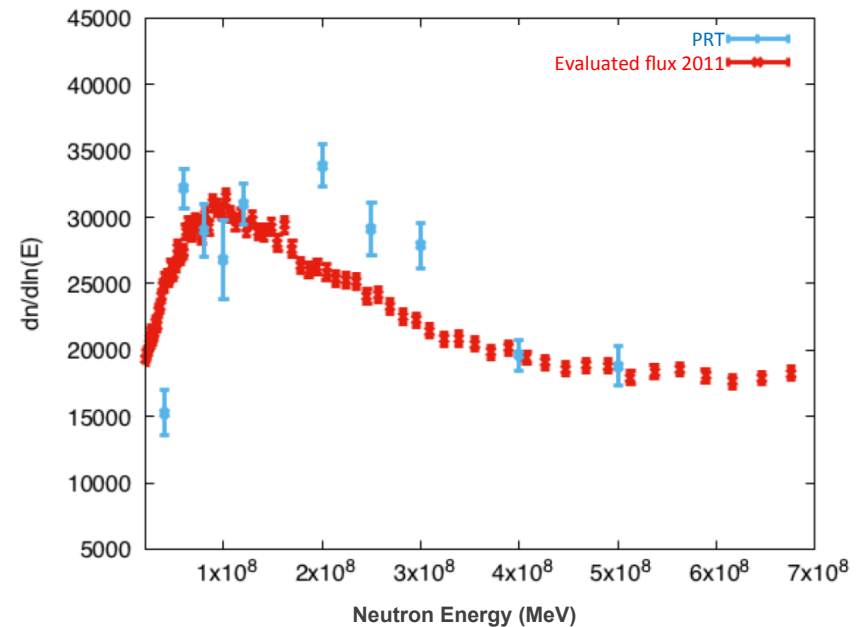
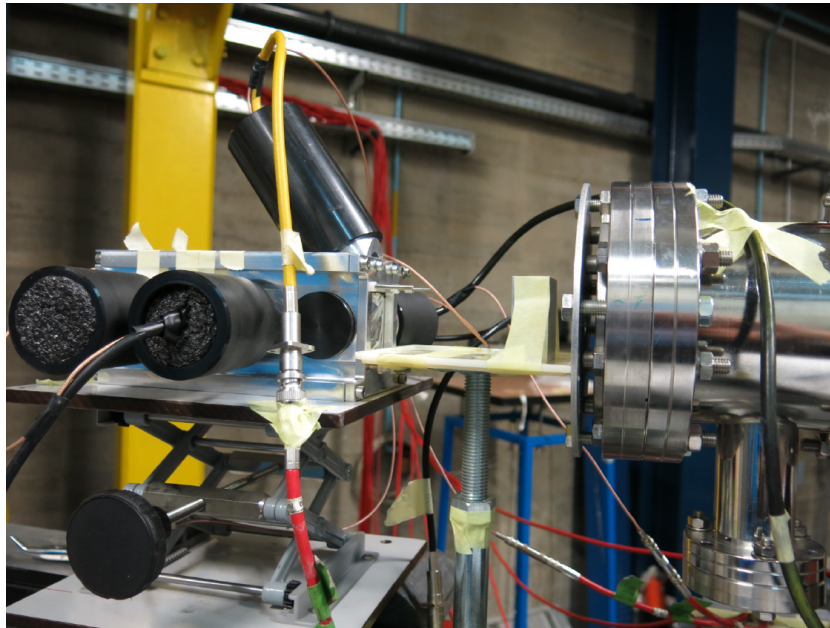
Energy spectra
Scint.4 @ $E_n \approx 200$ MeV

$\Delta E - E$ plot
 $E_n = 150 - 250$ MeV



A very preliminary flux reconstruction with the PRT prototype in the neutron energy range 10MeV – 500 MeV

(a few days test beam)



Alignment of the PRT to the neutron beam was not perfect

The amount of material in the PRT mechanics must be reduced, in order to cut down the background events

$$\varphi_n = \frac{C}{n\Omega\varepsilon\sigma}$$

C: counts
 n: sample areal density
 Ω: solid angle
 ε: efficiency
 σ: (n,p) cross section

- The Proton Recoil Telescope allows to distinguish the recoil protons from γ rays and heavier particles, in a wide neutron energy range

BUT

- A new mechanical setup and an optimized shielding for silicon detectors are needed, in order to suppress the background events

A final test of the PRT is foreseen at CERN in september 2017 with a renewed telescope sample, to finalize its configuration.

The measurement of the $^{235}\text{U}(n,f)$ cross section at n_TOF @CERN is planned for 2018.

L. Audouin¹, M. Barbagallo², N. Colonna², L. Cosentino², M. Diakaki^{3,4}, I. Duran⁵, P. Finelli^{6,2}, P. Finocchiaro², J. Heyse⁷, S. Lo Meo^{8,2}, C. Massimi^{2,6}, P.F. Mastinu², P.M. Milazzo², F. Mingrone³, A. Musumarra^{2,9}, R. Nolte¹⁰, C. Paradela⁷, D. Radeck¹⁰, P. Schillebeeckx⁷, L. Tassan-Got¹, G. Vannini^{2,6}, A. Ventura²

¹ CNRS-IN2P3, Univ. Paris-Sud, Univ. Paris-Saclay, Orsay, France

² INFN - sezioni di Bari, Bologna, LNL, LNS and Trieste, Italy

³ CERN, Switzerland

⁴ National Technical University of Athens, Greece

⁵ Universidad de Santiago de Compostela, Spain

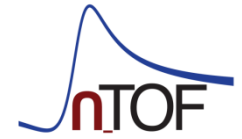
⁶ Dipartimento di Fisica e Astronomia, University of Bologna, Italy

⁷ European Commission, Joint Research Centre - Geel, Belgium

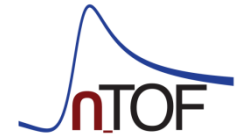
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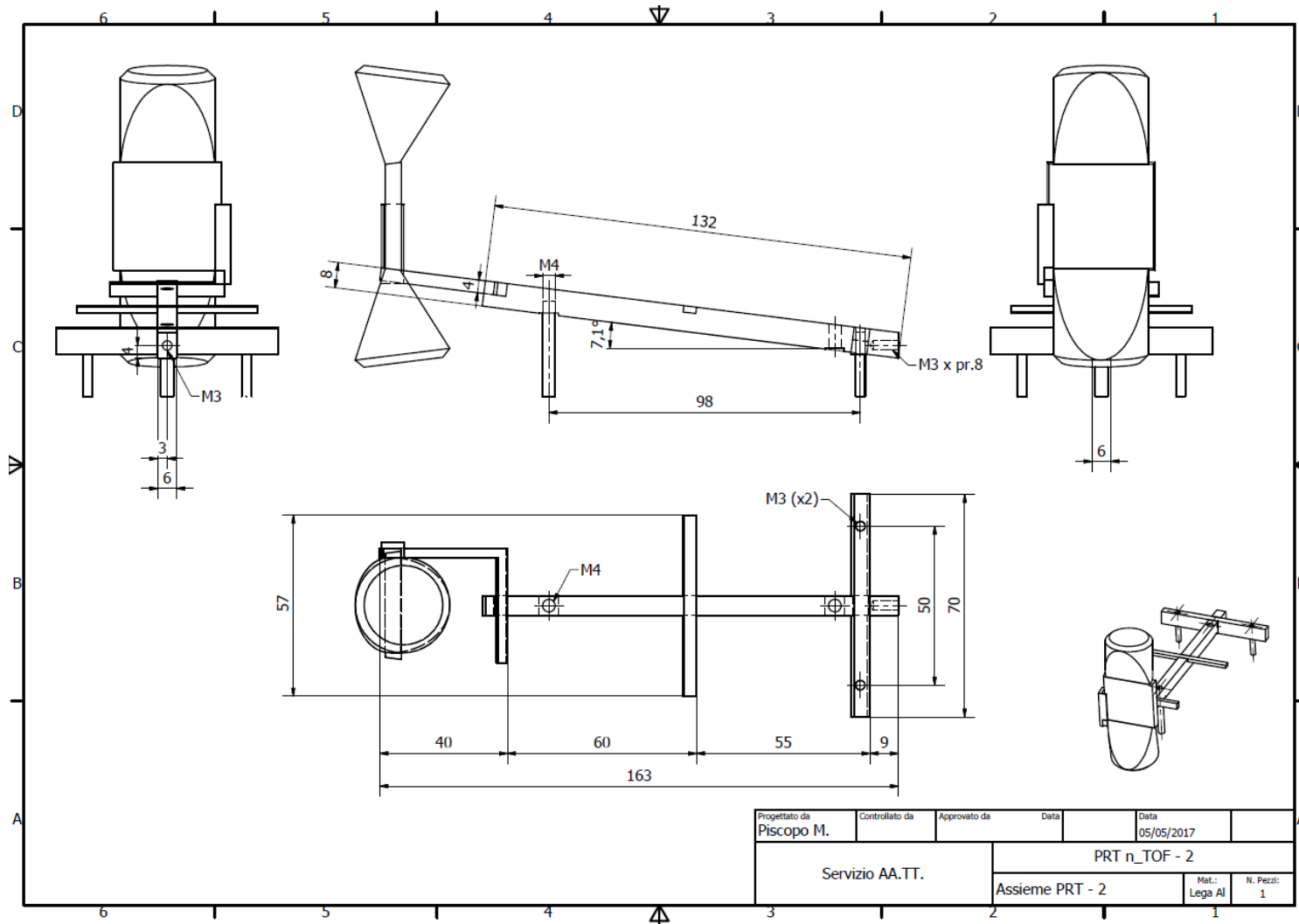
¹⁰ Physikalisch-Technische Bundesanstalt, Braunschweig, Germany

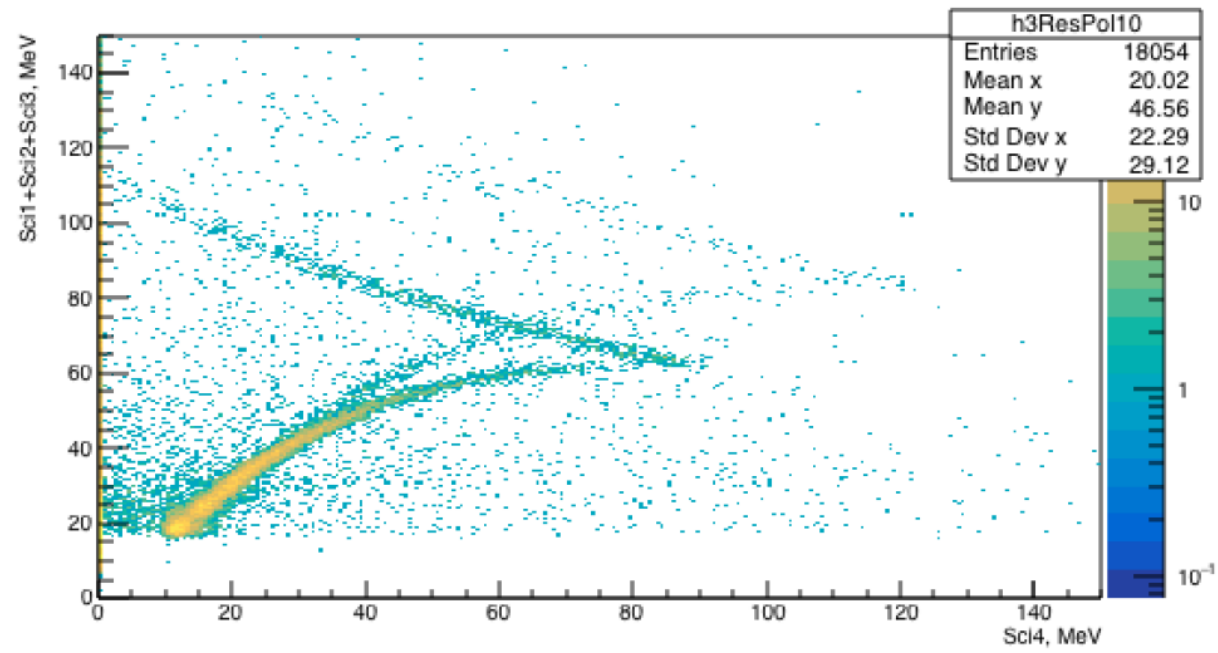


Thanks for your Attention



Backup Slides





900 MeV

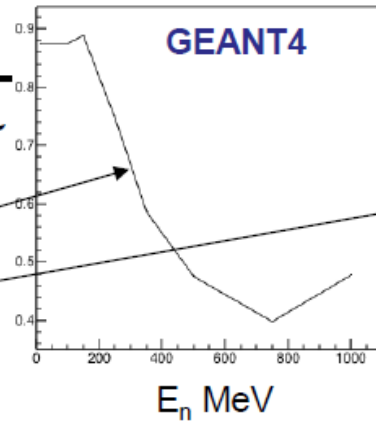
10-mm Polyethylene

$$\varphi_n = \frac{C}{n\Omega\varepsilon\sigma}$$

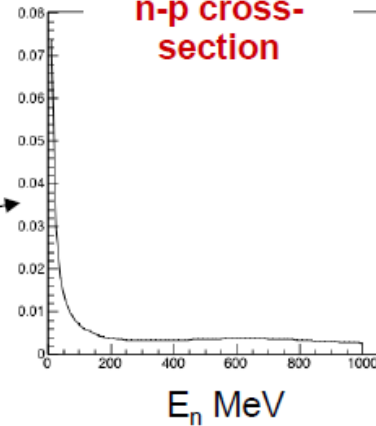
Sample
Areal density

Solid
angle

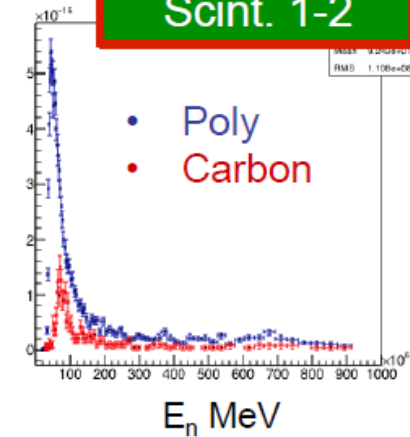
Efficiency



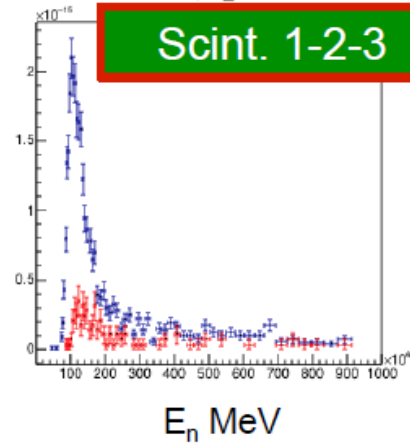
n-p cross-section



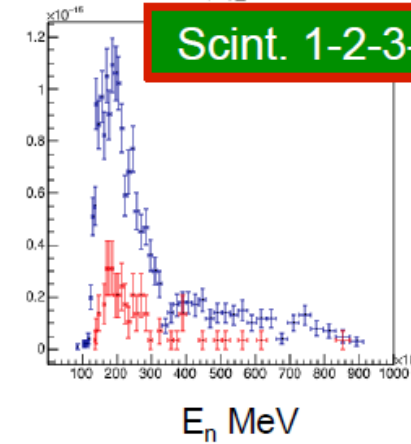
Scint. 1-2



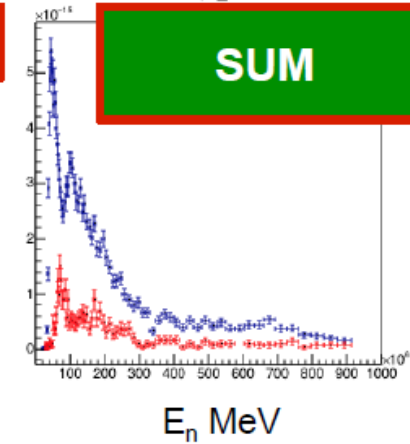
Scint. 1-2-3



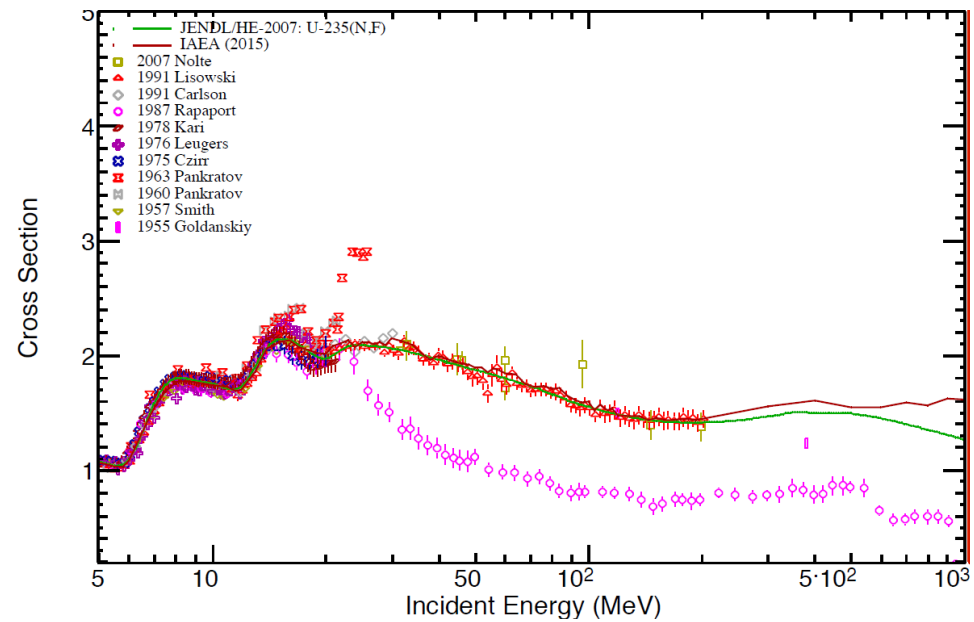
Scint. 1-2-3-4



SUM



Courtesy of Cristian Massimi



Theoretical estimates using the $^{235}\text{U}(p,f)$ reaction above 600 MeV as guidance.

Proposal to the ISOLDE and Neutron Time-of-Flight Committee

Measurement of the $^{235}\text{U}(n,f)$ cross section relative to n-p scattering up to 1 GeV

May 06, 2017

L. Audouin¹, M. Barbagallo², N. Colonna², L. Cosentino², M. Diakaki^{3,4}, I. Duran⁵, P. Finelli^{6,2}, P. Finocchiaro², J. Heyse⁷, S. Lo Meo^{8,2}, C. Massimi^{2,6}, P.F. Mastinu², P.M. Milazzo², F. Mingrone³, A. Musumarra^{2,9}, R. Nolte¹⁰, C. Paradela⁷, D. Radeck¹⁰, P. Schillebeeckx⁷, L. Tassan-Got¹, G. Vannini^{2,6}, A. Ventura²

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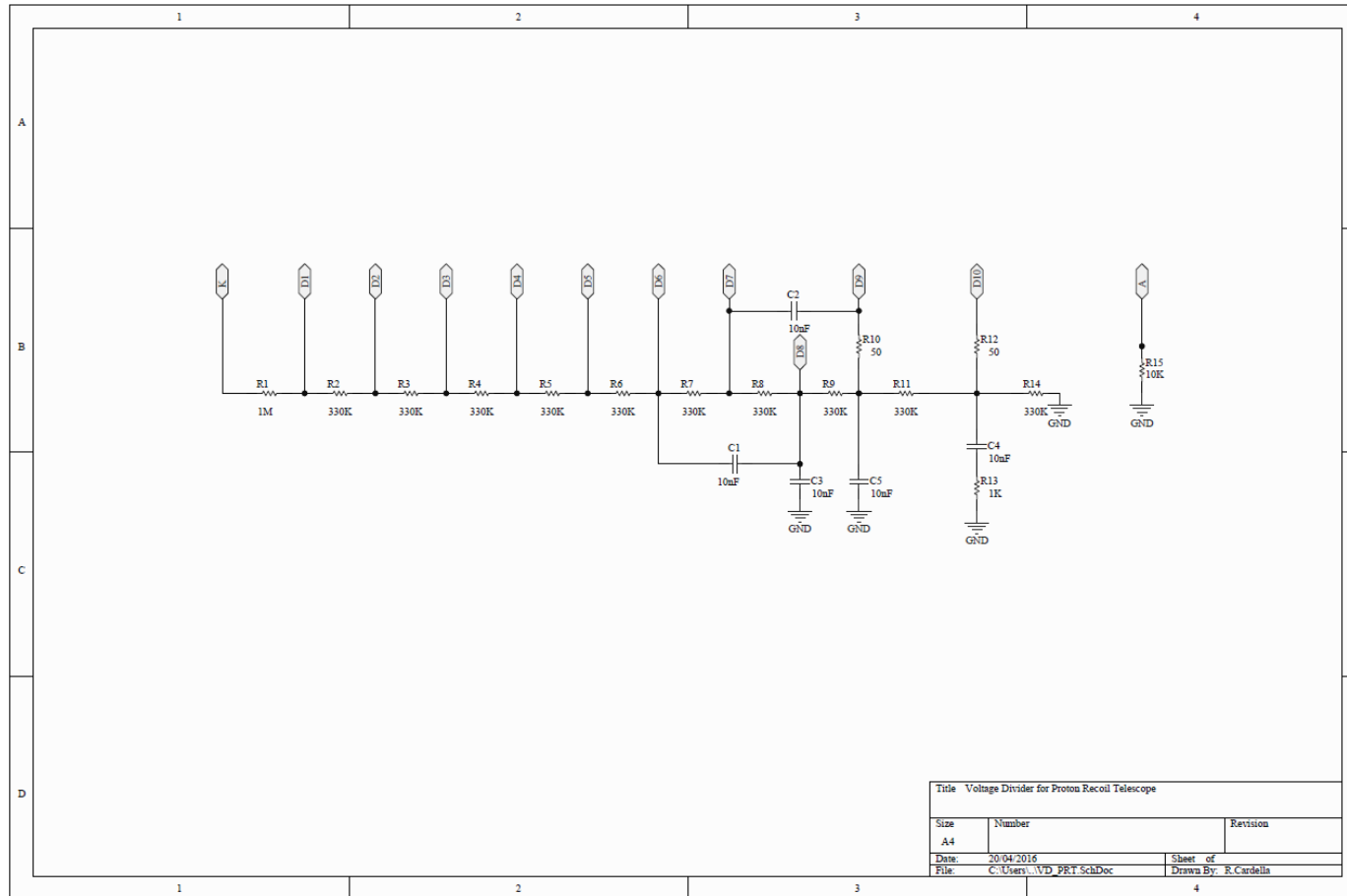
⁷ European Commission, Joint Research Centre - Geel, Belgium

⁸ ENEA - Bologna, Italy

⁹ Dipartimento di Fisica e Astronomia, University of Catania, Italy

¹⁰ Physikalisch-Technische Bundesanstalt, Braunschweig, Germany

Spokespersons: C. Massimi (massimi@bo.infn.it), R. Nolte (ralf.nolte@ptb.de) and
L. Cosentino (cosentino@lns.infn.it)
Technical coordinator: O. Aberle (oliver.aberle@cern.ch)



Analysis

- **Experimental data:**
 - $\Theta \approx (23 \pm 2)^\circ$
 - $\Delta E_1 = 2 \text{ mm}$, 5 mm EJ228, $\Delta E_2 = 5 \text{ mm}$ EJ228, $E = 80 \text{ mm}$ EJ204
 - No particle separation

- **Comparison with MC simulations (MCNPX 2.7):**
 - $E_n < 150 \text{ MeV}$: LA150 library for n and p
 - $E_n > 150 \text{ MeV}$: default INC model
 - Tracking of p and d
 - Non-linear light output simulated
 - $\Delta L/L = 10\%$ for E-detector

Courtesy of Ralf Nolte